



MAGNETOHYDRODYNAMICS EFFECTS ON STEADY NATURAL CONVECTION COUETTE FLOW OF HEAT GENERATING/ABSORBING FLUID IN A VERTICAL CHANNEL WITH VISCOUS DISSIPATION

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ABSTRACT

This study investigates the impact of magnetohydrodynamics (MHD) on steady natural convection Couette flow of a heat generating/absorbing fluid in a vertical channel in the presence of viscous dissipation. The dimensionless governing equations for momentum and energy were analytically solved using the Homotopy perturbation method. The influence of various physical parameters on the flow behavior is illustrated through graphical results. Moreover, the effects of these parameters on skin friction, rate of heat transfer, mass flux and mean temperature are analyzed and presented numerically in tabular form. Findings from the study show that, the magnetic field intensity on the velocity and temperature distributions of the fluid. It is worth highlighting that increasing the permeability enhances the fluid velocity and also the temperature near the heated plate rises, while a steeper temperature gradient is observed across the channel. Increasing the Eckert number significantly enhances both fluid velocity and temperature. The findings of this study are important due to its application in magnetohydrodynamics Power Generation, Metallurgy and Materials Processing, Cooling Systems in Nuclear Reactors, Electromagnetic Pumps, Plasma Propulsion and Aerospace and Food and Chemical Industries.

Keywords: Magnetic field, Heat generation/absorption, Homotopy perturbation method, Permeability, Viscous dissipation

INTRODUCTION

The study of magnetohydrodynamic (MHD) steady natural convection Couette flow with heat transfer in a vertical parallel channel has garnered significant attention in recent literature. These types of flows are prevalent in various industrial processes and natural phenomena, including geothermal energy extraction, nuclear waste management, groundwater flow, industrial and agricultural water systems, oil recovery, thermal insulation, pollutant transport in aquifers, cooling of electronic devices, packed bed reactors, and food processing. Numerous researchers have explored MHD flow in channels. Razaa et al. (2023) observed that fluid temperature rises with increased viscous dissipation. Jha and Aina (2017) found that accounting for the induced magnetic field leads to higher velocity profiles compared to when it is neglected. Omokhuale and Dange (2023) reported that an increase in magnetic field strength hinders fluid motion due to the Lorentz force. Ajibade and Tafida (2019a) emphasized the critical role of viscous dissipation in natural convection and recommended its inclusion in future studies. Akbar et al. (2022) demonstrated that fluid energy grows substantially with higher viscous dissipation. Waini et al. (2023) showed that a higher Eckert number reduces the heat transfer rate. Ajibade and Tafida (2019b) also noted that fluid velocity increases near the heated surface with a rise in the heat generation parameter. Hassan et al. (2022) concluded that both heat generation and absorption reduce the momentum boundary layer thickness. Usman (2024) found that stronger magnetic parameters enhance the Lorentz force, which opposes fluid motion within the boundary layer. Oni and Jha (2019) discovered that heat generation increases, while heat absorption decreases, both temperature and velocity distributions.

Tafida *et al.* (2023) observed that fluid velocity decreases as the magnetic field parameter increases. Ramesh (2018) noted that fluid temperature rises consistently with increasing viscous dissipation. Eswaran and Kumar (2025) emphasized flow velocity, while an increase in the magnetic parameter reduces it. Similarly, Ajibade *et al.* (2024) observed that

that understanding the influence of the magnetic parameter on fluid flow can contribute to better system design and optimization. Alam et al. (2018) found that increasing the magnetic field parameter results in lower velocity profiles but higher temperature profiles. Similarly, Hamza et al. (2024) reported that increasing the magnetic field strength leads to a reduction in fluid velocity. Jha and Samaila (2022) concluded that heat generation acts as a thermal resistance, impeding heat transfer from the left plate to the right plate. Tafida et al. (2024) observed that fluid velocity increases near the heated plate but decreases near the cold plate with greater viscous dissipation. In contrast, Kaita et al. (2024) found that fluid velocity increases with a rise in the magnetic field parameter. Zigta (2022) concluded that the temperature profile of a micropolar fluid decreases with an increase in the Prandtl number. Srisailam et al. (2023) demonstrated that while higher magnetic field values elevate the energy profile, they concurrently reduce the velocity profile.

Ajibade and Tafida (2020) found that both fluid temperature and velocity increase with higher heat generation, while they decrease with greater heat absorption. Mishra et al. (2023) observed that fluid velocity rises with increasing values of the magnetic field parameter. Kumar et al. (2023) reported that a rise in the Grashof number enhances buoyancy forces, which in turn intensifies fluid flow under heat generation and absorption conditions. Oni and Jha (2023) concluded that both skin friction and the rate of heat transfer are sensitive to variations in the heat generation/absorption parameter. Tafida and Tajuddeen (2024) emphasized that the Homotopy Perturbation Method provides a straightforward and effective approach for solving coupled and nonlinear differential equations. Nabwey et al. (2024) showed that magnetic permeability also influences the temperature distribution within the material. Shabiha and Tamizharasi (2025) found that fluid velocity decreases as the magnetic field strength increases. Sekhar (2023) noted that higher values of the permeability and heat generation parameters enhance fluid greater material permeability results in increased fluid velocity. Chillingo et al. (2024) observed that an increase in the permeability parameter leads to a gradual reduction in the fluid velocity profile.

Gouder et al. (2022) emphasized the Homotopy Perturbation Method (HPM) as a promising approach for wave equations. Similarly, Jassim and Mohammed (2021) found the method to be straightforward to implement while delivering accurate results. Sobamowo (2023) demonstrated Homotopy perturbation method capability to determine complex roots of nonlinear equations, while Khaleghizadeh (2022) affirmed its power and efficiency in tackling nonlinear problems. Farhood and Mohammed (2023) further noted that HPM is highly effective and enhances the rapid convergence of solutions through its simplicity.

In light of the reviewed literature, it is evident that further investigation into steady MHD natural convection flow and heat transfer in a vertical channel, particularly involving a heat-generating or absorbing fluid with viscous dissipation, is of considerable importance. This study aims to examine the effects of magnetohydrodynamics on steady natural convection Couette flow of such a fluid within a vertical channel, accounting for the influence of viscous dissipation. The governing equations are analytically solved using the Homotopy Perturbation Method. Graphical representations are provided to illustrate the impact of key physical parameters, while results for skin friction, heat transfer rate, mass flux, and mean temperature are presented in tabular form. The present investigation could be utilized in the food processing and preservation industries, as well as in fuel drilling and power generation sectors.

MATERIALS AND METHODS

Consider a steady fully developed laminar flow of a magnetohydrodynamic (MHD), viscous, incompressible and electrically conducting fluid, influenced by heat generation or absorption, viscous dissipation and a transverse magnetic field. One boundary plate remains stationary at ambient temperature while the opposing plate is heated and moves at a constant velocity U. The fluid is assumed to have constant physical properties throughout the flow domain.



Figure 1: Diagrammatic representation of the problem

The dimensional governing equations for momentum and energy are derived under the framework of the Boussinesq approximation and are expressed as follows:

$$v \frac{d^2 u^*}{dy^{*2}} - \left(\frac{\sigma B_0^2}{\rho} + \frac{v}{k^*}\right) u^* + g\beta(T^* - T_0) = 0,$$
(1)
$$k \frac{d^2 T^*}{dy^{*2}} - Q_0 (T^* - T_0) + \mu \left(\frac{du^*}{dy^*}\right)^2 = 0,$$
(2)

Here ν denotes the kinematic viscosity, u^* indicates the dimensional velocity, y^* refers to the dimensional distance, B_0 signifies the magnetic induction, β stands for the coefficient of thermal expansion, T^* represents the fluid temperature (in dimensional form), T_0 refers to the ambient temperature, k denotes the thermal conductivity, Q_0 is the coefficient of heat generation or absorption, ρ represents the fluid density, μ is the viscosity coefficient and g represents the acceleration due to gravity.

The system is subject to the following boundary conditions: $u^* = U$, $T^* = T$, at $v^* = 0$.

$$u^* = 0, T^* = T_0$$
 at $y^* = h$. (3)

Here, U denotes the velocity of the moving plate, T_w represents the temperature of the heated plate, and T_0 indicates the temperature of the cold plate. The dimensionless parameters used in equations (1) and (2) along with the boundary condition specified in equation (3), are defined as follows.

$$y = \frac{y^*}{h}, u = \frac{u^*}{U}, Ec = \frac{U_0}{cp(T_w - T_0)}, T = \frac{T^* - T_0}{T_w - T_0},$$

$$Pr = \frac{\nu}{\alpha}, Gr = \frac{g\beta h^2 (T_w - T_0)^2}{\nu U_0}, S = \frac{Q_0 h}{k}.$$
 (4)

Where y represents the non-dimensional temperature, u denotes the non-dimensional velocity, Pr is the Prandtl number, S refers to the non-dimensional heat generation/absorption parameter and Gr is the non-dimensional thermal Grashof number, respectively.

The governing momentum and energy equations (1) and (2) together with the boundary condition (3) are converted into their dimensionless form using the dimensionless equation (4) as shown below:

$$\frac{d^2u}{dy^2} - \left(M + \frac{1}{K}\right)u + GrT = 0,$$
(5)

$$\frac{d^2T}{dy^2} + Ec Pr\left(\frac{du}{dy}\right)^2 - ST = 0, \tag{6}$$

The associated boundary conditions are as follows: u = 1, T = 1 at y = 0,

$$u = 0, T = 0$$
 at $y = 1.$ (7)

Homotopy Perturbation Method

Using a convex homotopy, the momentum and energy equations (5) and (6), together with the boundary condition (7), are solved via the Homotopy perturbation method as outlined below:

$$H(u,p) = (1-p)\left(\frac{d^2u}{dy^2}\right) + p\left(\frac{d^2u}{dy^2} - \left(M + \frac{1}{K}\right)u + GrT\right) = 0,$$
(8)

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$$H(T,p) = (1-p)\left(\frac{d^2T}{dy^2}\right) + p\left(\frac{d^2T}{dy^2} + Ec Pr\left(\frac{du}{dy}\right)^2 - ST\right) = 0.$$
(9)

Accordingly, when no initial approximation is provided, the equation assumes the following form:

The solutions of equations (5) and (6) are expressed in the following form:

$$u = u_0 + pu_1 + p^2 u_2 + p^3 u_3 + \dots,$$
(12)

$$T = T_0 + pT_1 + p^2 T_2 + p^3 T_3 + \dots,$$
(13)

Substituting equation (12) into equation (10) and equation (13) into equation (11) yields the following results

$$\frac{d^{2}u_{0}}{dy^{2}} + p\frac{d^{2}u_{1}}{dy^{2}} + p^{2}\frac{d^{2}u_{2}}{dy^{2}} + \dots = p\left(\left(M + \frac{1}{\kappa}\right)u_{0} - GrT_{0}\right) + p^{2}\left(\left(M + \frac{1}{\kappa}\right)u_{1} - GrT_{1}\right) + p^{3}\left(\left(M + \frac{1}{\kappa}\right)u_{2} - GrT_{2}\right) + \dots \right)$$

$$\frac{d^{2}T_{0}}{dy^{2}} + p\frac{d^{2}T_{1}}{dy^{2}} + p^{2}\frac{d^{2}T_{2}}{dy^{2}} + \dots = p\left(-Ec\,Pr\left(\frac{du_{0}}{dy}\right)^{2} + ST_{0}\right) + p^{2}\left(-2Ec\,Pr\frac{du_{0}}{dy}\cdot\frac{du_{1}}{dy} + ST_{1}\right) + p^{3}\left(-2Ec\,Pr\frac{du_{0}}{dy}\cdot\frac{du_{2}}{dy} - Ec\,Pr\left(\frac{du_{1}}{dy}\right)^{2} + ST_{2}\right) + \dots$$
(15)

Equating the coefficients of the corresponding terms p^0 , p^1 , p^2 and p^3 in equations (14) and (15) gives: $d^2 u_0$

$$p^{0}: \frac{dy^{2}}{dy^{2}} = 0, \tag{16}$$

$$p^{0}: \frac{d^{2}T_{0}}{dy^{2}} = 0, \tag{17}$$

$$p^{1} \cdot \frac{dy^{2}}{dy^{2}} = \left(M + \frac{1}{K}\right)u_{0} - GrT_{0},$$
(18)

$$p^{1}:\frac{d^{2}T_{1}}{dy^{2}} = -Ec Pr\left(\frac{du_{0}}{dy}\right)^{2} + ST_{0},$$
(19)

$$p^{2}:\frac{d^{2}u_{2}}{dy^{2}} = \left(M + \frac{1}{\kappa}\right)u_{1} - GrT_{1},$$
(20)

$$p^{2}:\frac{d_{12}}{dy^{2}} = -2Ec Pr\left(\frac{du_{0}}{dy},\frac{du_{1}}{dy}\right) + ST_{1},$$

$$n^{3}:\frac{d^{2}u_{3}}{dy} = \left(M + \frac{1}{2}\right)u_{1} - CrT_{2}$$
(21)

$$p^{3} : \frac{dy^{2}}{dy^{2}} = -2Ec Pr\left(\frac{du_{0}}{dy} \cdot \frac{du_{2}}{dy}\right) - Ec Pr\left(\frac{du_{1}}{dy}\right)^{2} + ST_{2}.$$
(23)

The boundary condition given in equation (7) is reformulated as:

$$\begin{array}{l} u_{0}(0) = 1, u_{1}(0) = u_{2}(0) = u_{3}(0) = \dots = 0, \\ u_{0}(1) = u_{1}(1) = u_{2}(1) = u_{3}(1) = \dots = 0, \\ T_{0}(0) = 1, T_{1}(0) = T_{2}(0) = T_{3}(0) = \dots = 0, \\ T_{0}(1) = T_{1}(1) = T_{2}(1) = T_{2}(1) = \dots = 0. \end{array}$$

$$(24)$$

By applying equations (16) and (17) in conjunction with the boundary conditions $u_0(0) = 1$ and $u_0(1) = 0$, $T_0(0) = 1$ and $T_0(1) = 0$ performing the necessary computations, equations (25) and (26) are derived as follows:

$$u_0 = A_1 y + A_2, (25)T_0 = B_1 y + B_2. (26)$$

By applying equations (18) and (19) in conjunction with the boundary conditions $u_1(0) = 0$ and $u_1(1) = 0$, $T_1(0) = 0$ 0 and $T_1(1) = 0$ performing the necessary computations, equations (27) and (28) are derived as follows:

$$u_{1} = \left(M + \frac{1}{K}\right)\left(\frac{y^{2}}{2} - \frac{y^{3}}{6}\right) - Gr\left(\frac{y^{2}}{2} - \frac{y^{3}}{6}\right) + A_{3}y + A_{4},$$
(27)
$$T_{1} = -\frac{Ec Pr y^{2}}{2} + S\left(\frac{y^{2}}{2} - \frac{y^{3}}{6}\right) + B_{3}y + B_{4}.$$
(28)

By applying equations (20) and (21) in conjunction with the boundary conditions $u_2(0) = 0$ and $u_2(1) = 0$, $T_2(0) = 0$ 0 and $T_2(1) = 0$ performing the necessary computations, equations (29) and (30) as follows:

$$\begin{split} u_2 &= \left(M + \frac{1}{\kappa}\right)^2 \left(\frac{y^4}{24} - \frac{y^5}{120}\right) - Gr\left(M + \frac{1}{\kappa}\right) \left(\frac{y^4}{24} - \frac{y^5}{120}\right) + \\ \left(M + \frac{1}{\kappa}\right) \frac{A_3 y^3}{6} + \frac{Ec \, Pr \, Gr y^4}{24} - Gr S\left(\frac{y^4}{24} - \frac{y^5}{120}\right) - \frac{Gr B_3 y^3}{6} + \\ A_5 y + A_6, & (29) \\ T_2 &= Ec \, Pr \, Gr\left(M + \frac{1}{\kappa}\right) \left(\frac{y^3}{3} - \frac{y^4}{12}\right) - Ec \, Pr \, Gr \left(\frac{y^3}{3} - \frac{y^4}{12}\right) + \\ Ec \, Pr \, A_3 \, y^2 - \frac{Ec \, Pr \, Sy^4}{24} + S^2 \left(\frac{y^4}{24} - \frac{y^5}{120}\right) + \frac{SB_3 y^3}{6} + B_5 y + \\ B_6, & (30) \\ \text{where,} \\ A_1 &= B_1 = -1, \\ A_2 &= B_2 = 1, \\ A_3 &= \frac{Gr}{3} - \frac{1}{3} \left(M + \frac{1}{\kappa}\right), \ A_4 = 0, \\ B_3 &= \frac{Ec Pr \, Gr}{2\frac{S}{3}} B_4 = 0, \\ A_5 &= -\frac{1}{30} \left(M + \frac{1}{\kappa}\right)^2 + \frac{Gr}{30} \left(M + \frac{1}{\kappa}\right) - \frac{A_3}{6} \left(M + \frac{1}{\kappa}\right) - \\ \frac{Ec \, Pr \, Gr}{4} + \frac{GrS}{4} + \frac{GrB_3}{4} \end{split}$$

$$\frac{1}{24} + \frac{1}{30} + \frac{1}{6},$$

$$B_5 = \frac{Ec\,Pr\,Gr}{4} - \frac{EcPr}{4} \left(M + \frac{1}{K}\right) - EcPr\,A_3 + \frac{Ec\,Pr\,S}{24} - \frac{S^2}{30} - \frac{SB_3}{6}$$

 $A_6 = B_6 = 0.$

Equations (25) - (30) give the approximate solutions for velocity and temperature as:

$$u = u_0 + u_1 + u_2 + \dots,$$
(31)

$$T = T_0 + T_1 + T_2 + \dots$$
(32)

The skin friction at the channel fluid boundary surfaces is determined by differentiating equation (31) with respect to y and evaluated at the boundaries y = 0 and y = 1 as follows:

$$\tau_{0} = \frac{du}{dy}\Big|_{y=0} = -1 + A_{3} + A_{5}, \qquad (33)$$

$$\tau_{1} = \frac{du}{dy}\Big|_{y=1} = -1 + \frac{1}{2}\left(M + \frac{1}{K}\right) - \frac{Gr}{2} + A_{3} + \frac{1}{8}\left(M + \frac{1}{K}\right)^{2} - \frac{Gr}{8}\left(M + \frac{1}{K}\right) + \frac{A_{3}}{2}\left(M + \frac{1}{K}\right) + \frac{EcPr}{6\frac{GrSGrB_{3}}{8}} \qquad (34)$$

The heat transfer rate at the channel fluid boundary surfaces is determined by differentiating equation (32) with respect to y and evaluated at the boundaries y = 0 and y = 1 as follows: $Nu_{0} = \frac{dT}{dy}\Big|_{y=0} = -1 + B_{3} + B_{5}, \qquad (35)$ $Nu_{1} = \frac{dT}{dy}\Big|_{y=1} = -1 - Ec Pr + \frac{SB_{3}}{2} + \frac{2EcPr}{3\left(M + \frac{1}{K}\right)^{\frac{2EcPrGr}{3}}}$ (35) $+2Ec Pr A_3 - \frac{Ec Pr S}{6} + \frac{S^2}{8} + \frac{SB_3}{2} + B_5.$ (36) Using the expression for mass flux $Q_m = \int_0^1 u(y) dy$, we

derive:

$$\begin{split} Q_m &= \frac{1}{2} + \frac{1}{2} \left(M + \frac{1}{K} \right) - \frac{Gr}{8} + \frac{A_3}{2} + \frac{1}{144} \left(M + \frac{1}{K} \right)^2 - \frac{Gr}{144} \left(M + \frac{1}{K} \right) \\ &+ \frac{A_3}{24} \left(M + \frac{1}{K} \right) + \frac{Ec\,Pr\,Gr}{120} - \frac{GrS}{144} - \frac{GrB_3}{24} + \frac{A_5}{2}, \quad (37) \end{split}$$

In a similar manner, the mean temperature is expressed as:
$$\theta_m &= -\frac{1}{3} - \frac{7EcPr}{24} + \frac{13S}{60} + \frac{5B_3}{6} + \frac{1}{30} \left(M + \frac{1}{K} \right) - \frac{13EcPr}{360} \left(M + \frac{1}{K} \right) \\ &+ \frac{1}{K} \right) \\ &\frac{A_3}{6} + \frac{S}{65} \left(M + \frac{1}{K} \right) + \frac{B_3}{8} \left(M + \frac{1}{K} \right) - \frac{Gr}{30} + \frac{13EcPr\,Gr}{360} + \frac{GrS}{64} - \frac{11GrB_3}{120} - \frac{EcPr\,A_3}{8} + \frac{11AS}{120} + \frac{A_3B_3}{3}. \quad (38) \end{split}$$

RESULTS AND DISCUSSION

This study investigates a magnetohydrodynamic natural convection model characterized by several fundamental physical parameters, such as viscous dissipation, Prandtl number, Grashof number, heat generation or absorption, magnetic field and permeability. The effects of these parameters on velocity and temperature distributions are depicted through graphical illustrations. Furthermore, numerical values for skin friction, heat transfer rate, mass flux and mean temperature are presented in tabular form for comprehensive analysis

 v^5 \



Figure 2: Velocity profile for different values of Ec when Gr = 6.0, Pr = 0.71, M = 0.4, S = 1.0 and K = 0.4.

As depicted in Figures 2 and 3, increasing the Eckert number significantly enhances both fluid velocity and temperature, primarily as a result of intensified viscous dissipation. This phenomenon mitigates the damping effects imposed by the magnetic field and exemplifies the behavior of thermo-fluid systems where internal energy transformation plays a crucial role in governing flow characteristics. The resulting



Figure 4: Velocity profile for different values of Pr when Ec = 0.4, Gr = 6.0, M = 0.4, S = 1.0 and K = 0.4.

Figures 4 and 5 illustrate the velocity and temperature profiles, respectively, for different values of the Prandtl number. From Figure 4, it is evident that an increase in the Prandtl number leads to a slight rise in the velocity profile, with the peak shifting upward before gradually declining toward the wall. This trend can be attributed to the reduced thermal diffusivity associated with higher Prandtl numbers, which allows heat to remain in the fluid for a longer duration,



when Gr = 6.0, Pr = 0.71, M = 0.4, S = 1.0 and K = 0.4

temperature rise amplifies buoyant forces, thereby promoting greater fluid acceleration. Additionally, a higher Eckert number augments internal heat generation, leading to elevated temperature levels throughout the boundary layer. In essence, increased viscous dissipation facilitates a more efficient conversion of kinetic energy into thermal energy.



Figure 5: Temperature profile for different values of Pr when Ec = 0.4, Gr = 6.0, M = 0.4, S = 1.0 and K = 0.4.

thereby intensifying the buoyancy effect and enhancing fluid motion. In Figure 5, it is observed that as the Prandtl number increases, the temperature near the heated wall becomes higher, while the temperature drops more sharply across the channel. This indicates that higher Prandtl numbers result in steeper temperature gradients due to the lower thermal conductivity of the fluid.



Figure 6: Velocity profile for different values of Gr when Ec = 0.4, Pr = 0.71, M = 0.4, S = 1.0 and K = 0.4.

Figure 6 illustrates the influence of the Grashof number on fluid velocity. It can be observed that an increase in the Grashof number leads to a notable rise in fluid velocity, attributed to the intensification of buoyancy driven forces. In practical applications where natural convection governs the flow, modifying parameters that affect the Grashof number, such as fluid properties or temperature gradients serves as a practical strategy to enhance both flow dynamics and heat transfer efficiency.



Figure 7: Temperature profile for different values of Gr when Ec = 0.4, Pr = 0.71, M = 0.4, S = 1.0 and K = 0.4.

Figure 7 depicts the corresponding effect on the temperature profile. As the Grashof number increases, the temperature field becomes more diffused, indicating stronger convective









Figures 8 and 9 illustrate the impact of heat generation and absorption on the velocity and temperature profiles. As heat generation increases, both the fluid temperature and velocity rise. Conversely, increased heat absorption results in a reduction of these parameters. This decline is attributed to the drop in fluid temperature, which leads to a thinner thermal boundary layer and a corresponding decrease in fluid velocity.



Figures 10 and 11 depict the effect of magnetic field intensity on the velocity and temperature distributions of the fluid. Figure 10 demonstrates that fluid velocity increases as the magnetic field strength rises. This behavior aligns with physical reasoning, as a stronger magnetic field amplifies the



thicker thermal boundary layer and enhancing the fluid's convective flow. In addition, increased heat generation boosts both temperature and fluid motion, thereby amplifying convective heat transfer mechanisms. M = 0.4M = 0.6M = 0.8M = 1.2

On the other hand, higher heat generation elevates the

temperature and velocity distributions, contributing to a



when Ec = 0.4, Pr = 0.71, Gr = 6.0, S = 1.0 and K = 0.4.

Lorentz force, which resists fluid motion and consequently elevates the velocity. Furthermore, an increase in magnetic field strength leads to a rise in fluid temperature near the heated plate, while a reduction in temperature is observed near the cold plate.



The influence of permeability parameter on velocity and temperature are depicted in Figures 12 and 13 illustrate the effect of the permeability parameter Kon the fluid velocity and temperature profiles. As shown in Figure 12, increasing the permeability enhances the fluid velocity, as higher permeability facilitates freer movement of fluid particles,

thereby promoting greater flow. Figure 13 demonstrates that with increasing permeability, the temperature near the heated plate rises, while a steeper temperature gradient is observed across the channel.

Table 1: Estimated Numerical V	Values of skin Frict	tion at the Plate $v =$	= 0 and $y = 1$
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	(Ec = 0.2, K = 0.4, M = 1.0)		(Ec = 0.2, K = 0.4, M = 2.0)		(Ec = 0.4, K = 0.8, M = 2.0)	
S	$ au_0$	$ au_1$	$ au_0$	$ au_1$	$ au_0$	$ au_1$
-1.0	0.43650	0.24321	0.41967	0.22034	1.34662	1.19890
-0.5	0.46704	0.25018	0.44281	0.23551	1.36414	1.21739
0	0.47132	0.27664	0.46300	0.25920	1.39001	1.40818
0.5	0.50167	0.29541	0.48421	0.27684	1.41238	1.69200
1.0	0.53103	0.31282	0.50132	0.29965	1.42998	1.80412

Table 1 illustrates the variation in skin friction between the fluid and the bounding plates. It is observed that an increase in magnetic field strength leads to a reduction in skin friction on both plates. This behavior is consistent with physical expectations, as a stronger magnetic field intensifies the Lorentz force, which acts against the fluid motion, thereby diminishing skin friction. Furthermore, the data indicate that skin friction increases with greater viscous dissipation and higher permeability of the medium. Lastly, skin friction is found to rise with enhanced heat absorption, while it decreases in response to increased heat generation.

Table 2: Estimated numerical values of rate of heat transfer at the plate $y = 0$ and $y = 0$	= 1	1
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c	(Ec = 0.2, K = 0.4, M = 1.0)		(Ec = 0.2, K = 0.4, M = 2.0)		(Ec = 0.4, K = 0.8, M = 2.0)	
3 -	Nu ₀	Nu ₁	Nu_0	Nu ₁	Nu ₀	Nu ₁
-1.0	1.44640	1.33314	1.43092	1.31114	1.57211	1.62013
-0.5	1.44042	1.35012	1.42901	1.32915	1.54013	1.64193
0	1.42125	1.38864	1.40311	1.33202	1.51071	1.67010
0.5	1.40307	1.39997	1.38121	1.35104	1.48216	1.69982
1.0	1.38201	1.40512	1.37426	1.37252	1.45813	1.72236

Table 2 highlights the rate of heat transfer between the fluid and the plates. The results indicate that an increase in magnetic field strength suppresses the heat transfer rate on both plates. In contrast, higher values of viscous dissipation and permeability lead to an enhancement in heat transfer. The data further reveal that heat absorption increases the heat transfer rate on the moving plate, whereas a reduction is observed on the stationary plate. On the other hand, increasing heat generation results in a decline in heat transfer on the moving plate, but an increase on the stationary plate. This contrasting behavior is attributed to the movement of high energy fluid particles from the moving plate toward the stationary plate, which raises the fluid temperature in the vicinity of the stationary surface.

 Table 3: Estimated numerical values of mass flux Q

S	(Ec = 0.2, K = 0.4, M = 1.0)	(Ec = 0.2, K = 0.4, M = 2.0)	(Ec = 0.4, K = 0.8, M = 2.0)
	Q	Q	Q
-1.0	0.47716	0.49241	0.53056
-0.5	0.49562	0.52087	0.55501
0	0.53281	0.55470	0.58109
0.5	0.55815	0.57805	0.60411
1.0	0.57611	0.59257	0.63055

Table 3 provides the numerical values of mass fluxQ. The results reveal that an increase in magnetic field strength leads to an enhancement in mass flux. Additionally, higher values of viscous dissipation and permeability are associated with

increased mass flux. Moreover, heat absorption is found to promote mass flux, while heat generation has the opposite effect causing a reduction.

Table 4: E	stimated n	umerical v	alues of	f mean	temperature 6	Э,
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S	(Ec = 0.2, K = 0.4, M = 1.0)	(Ec = 0.2, K = 0.4, M = 2.0)	(Ec = 0.4, K = 0.6, M = 2.0)	
	$\boldsymbol{ heta}_{m}$	$\boldsymbol{ heta}_{m}$	$\boldsymbol{\theta}_{\boldsymbol{m}}$	
-1.0	0.42105	0.42892	0.49164	
-0.5	0.37961	0.38861	0.44591	
0	0.35234	0.35912	0.41602	
0.5	0.32398	0.33239	0.39101	
1.0	0.31056	0.31786	0.36231	

Table 4 presents the numerical values for the mean temperature. It is evident that both viscous dissipation and the permeability parameter contribute to an increase in mean temperature, reflecting their role in converting mechanical

energy into thermal energy. Moreover, the mean temperature rises with an increase in heat generation, whereas the opposite effect is observed with heat absorption.

Validation

Fable 5: Validation o	present	problem and	that of A	jibade and	Tafida ((2019a))
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Ajibade and Tafida (2019a) <i>S Ec</i> = 0. 4, <i>Pr</i> = 0. 71, <i>Gr</i> = 8. 0, <i>y</i> = 0. 5.		Present Problem Ec = 0.4, Pr = 0.71, Gr = 8.0, y = 5.0, M = 0, K = 10		
	Velocity	Temperature	Velocity	Temperature
-1.0	1.06415	0.56833	1.06407	0.56827
-0.5	1.02906	0.53502	1.02914	0.53513
0.5	0.97528	0.47038	0.97516	0.47025
1.0	0.95398	0.44349	0.95385	0.44335

For numerical validation of the present problem, the computed results for velocity and temperature are compared

with the published work of Ajibade and Tafida (2019a), as presented in Table 5. A good agreement is observed when the

magnetic field is neglected and the permeability parameter is assumed to be sufficiently large.

CONCLUSION

This study presents a mathematical analysis and solution of steady MHD natural convection Couette flow of an incompressible fluid within vertical channels. The dimensionless governing momentum and energy equations are solved using the homotopy perturbation method. The influence of various dimensionless parameters on the flow characteristics is illustrated and analyzed through graphical representations. The results show excellent agreement with the findings of Ajibade and Tafida (2019a). The key conclusions of this investigation are summarized as follows:

- i. An increase in the Eckert number leads to a significant rise in both fluid velocity and temperature.
- ii. Increased heat generation results in elevated temperature and velocity profiles,
- iii. An increase in magnetic field strength enhances fluid velocity and raises the temperature near the heated plate,
- iv. Higher permeability increases fluid velocity and also elevates the temperature near the heated plate,
- v. An increase in magnetic field strength reduces the heat transfer rate on both plates,
- vi. Finally, magnetic field and permeability are essential fluid properties, and their incorporation into the fluid flow equations plays a key role in shaping the flow dynamics and thermodynamic characteristics within the channel

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